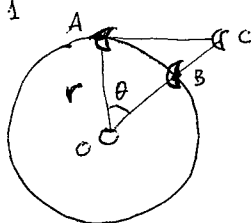


LAW OF UNIVERSAL GRAVITATION

Galileo ⁽⁺⁾ Rock falling ⁽⁺⁾ distance $d = \frac{gt^2}{2}$ for t sec's, $g = 9.81 \text{ m/s}^2 = a_0$ (accel of rock)

Newton's hypothesis 1 The moon behaves like a falling rock ⁽⁺⁾ (equilibrium) some acceleration $a = \frac{a_0}{r}$. After Δt sec's



$$\frac{a(\Delta t)^2}{2} \rightarrow BC = CO - OB = \frac{r}{\cos\theta} - r = r \frac{1 - \cos\theta}{\cos\theta}$$

$$= r \frac{1 - \cos \frac{2\pi}{T} \Delta t}{\cos \frac{2\pi}{T} \Delta t} = 2r \frac{\sin^2 \frac{\pi}{T} \Delta t}{\cos \frac{2\pi}{T} \Delta t}$$

$$\frac{\theta}{2\pi} = \frac{\Delta t}{T}$$

$$\Rightarrow a \sim 4r \left(\frac{\sin \frac{\pi}{T} \Delta t}{\Delta t} \right)^2 \sim 4r \frac{\pi^2}{T^2}$$

(cos 0 = 1)

Same for planets \searrow around Sun \swarrow

$$\frac{\sin ax}{x} \sim a$$

Kepler's 3rd law $o = \text{planet}$, $\star = \text{Sun}$

$$\frac{r_o^3}{T_o^2} = C_{\star}, \text{ act dep on } \star \text{ but not on } o$$

Cor $a_o = 4r_o \frac{\pi^2}{T_o^2} = 4 \frac{r_o^3}{r_o^2} \frac{\pi^2}{T_o^2} = \frac{4\pi^2 C_{\star}}{r_o^2}$

Newton's hyp 2 $a_g = \frac{4\pi^2 C_o}{r_g^2}$, $a_o = \frac{4\pi^2 C_o}{r_o^2}$

more generally for any 2 bodies 1 & 2
 $a_1 = 4\pi^2 C_2 / r_{12}^2$

sun	planets
Earth	Moon (or rock)

Numerical verif $9.81 \text{ m/s}^2 = a_o = \frac{4\pi^2 C_o}{r_o^2} = \frac{4\pi^2}{4\pi^2} \frac{a_g r_g^2}{r_o^2} = \frac{1}{r_o^2} \left(4r_g \frac{\pi^2}{T_g^2} \cdot r_g^2 \right) \sim 9.81$

(dir calc ⁽⁺⁾)

Def (Newton) A phys sys is a set of n bodies $\vec{x}_i: \mathbb{R} \rightarrow \mathbb{R}^3$ (trajectories), a_g $\vec{F}_{ij}: \mathbb{R} \rightarrow \mathbb{R}^3$ (forces)
& #'s $m_i \in \mathbb{R}$ (masses) s.t. $\vec{F}_{ij} + \vec{F}_{ji} = 0$ & $m_i \frac{d^2 \vec{x}_i}{dt^2} = \vec{F}_i$, $\vec{F}_i = \sum_j \vec{F}_{ij}$

Deduction of law of univ grav 2 bodies $\vec{F}_{ij} = |\vec{F}_{ij}| \Rightarrow F_{12} = F_{21}$
 $m_1 \frac{4\pi^2 C_2}{r_{12}} = m_1 a_1$, $m_2 a_2 = m_2 \frac{4\pi^2 C_1}{r_{12}}$

$$\Rightarrow \frac{C_1}{m_1} = \frac{C_2}{m_2} \text{ for } \forall 2 \text{ bodies in universe; call this } \frac{k}{4\pi^2}$$

$$\Rightarrow F_{12} = \frac{4\pi^2}{r_{12}^2} m_1 m_2 \frac{C_2}{m_2} = \frac{k m_1 m_2}{r_{12}^2}$$

⁽⁺⁾ thrown horizontally

⁽⁺⁾ a vertical ⁽⁺⁾

$r_o = \text{rad of Earth}$
 $r_g = \text{dis Earth to Moon}$
 $T_g = 28 \text{ days}$